## Lec #18: 3 OCT 2011

- **TODAY: Spectral Sequence Explained**
- What happens if we don't have TE?
- LTE Distribution Functions
- Line Formation v. Temperature
- NEXT: Atomic Physics
- The Bohr Atom
- Hydrogenic Ions
- Modifications required for >1 electron
- Atomic Structure, Selection Rules, Atomic "Terms"

## $\label{eq:transformation} \begin{array}{l} \hline TE \ Distribution \ Functions \\ \bullet \ Maxwell-Boltzmann \ Distribution: \\ N_v/N \ (T_k) \ dv = [m/2\pi kT_K]^{3/2} \ e^{-mv^2/2kT_K} \ 4\pi v^2 \ dv \\ \bullet \ Boltzmann \ Distribution: \\ N_b/N_a \ (T_x) = (g_b/g_a) \ e^{-(E_b-E_a)/kT_x} \\ \bullet \ Saha \ Equation: \\ N_{i+1}/N_i \ (T_i) = (2/n_e) \ (Z_{i+1}/Z_i) \ (2\pi m_e kT_i/h^2)^{3/2} \ e^{-\chi_i/kT_i} \\ \bullet \ Planck \ Function: \\ B_v(T_r) = \ 2hv^3/c^2 \ [e^{hv/KT_r} \ -1]^{-1} \\ \bullet \ \ Where \ does \ this \ e^{-\Delta E/kT} \ come \ from? \end{array}$

## **Probability and Statistical Physics**

• Probability =

(# of ways to achieve desired result) /
(# of ways to achieve any result)

- # of ways energy can be distributed among states with kT per degree of freedom increases exponentially with Energy (i.e. temperature)
- so # of states ~  $e^{\Delta E/kT}$
- postulate: all possible states are equally likely
- so P is # of ways to put energy in desired state divided by number of possible states ~ e<sup>-ΔE/kT</sup>











- use  $\chi_i = 13.6 \text{ eV}$ 











