

- **LAST TIME:** RA & Dec; Daily Motion of the Sky
 - Spherical Earth
 - Latitude and Longitude
 - Right Ascension and Declination
- **TODAY:** Annual and Long-Term Motions of Sky
 - Daily Motion of Sky from any latitude on Earth
 - Apparent Annual Motion of Sun and Nighttime Sky
 - Putting it all together
 - Long-Term Variations: Parallax, ...
- **FRIDAY:** Practice, Review, Digest Coordinate Systems
- **NEXT Week:** Cataloging Stellar Properties. I.
 - how do we measure position accurately/precisely?
 - what can we learn from it?

The Sky Viewed from Charleston

- motion depends on where in sky you are looking
- some apparent paths are east->west arcs
- some apparent paths are counterclockwise circles
- time from rise to set depends on declination:
 $0^h: < -57^\circ; < 12^h: -57 \text{ to } 0^\circ; > 12^h: 0-57^\circ; 24^h > 57^\circ$



(a) At middle northern latitudes

- NCP elevation = 33°
- CE from east through meridian elevated by 57° to west
- see all stars north of declination -57°



- all stars move in ccw circles around NCP
- stars with declination more than 90 minus your latitude (57 to 90 degrees for Charleston) are "CIRCUMPOLAR"
- for stars south of this, we only see a portion of their circular path, so they appear to move in an arc
- notice that stars farther from pole move farther in same amount of time

The Sky From Different Places on Earth

- everyone on same longitude sees same "time" (i.e. Sun same distance east or west of meridian)
- everyone on *same latitude* sees the *same part of the sky* (e.g. 33° north latitude sees from -57° to $+90^\circ$ of declination) at the *same local time*, but local time depends on longitude
- e.g., at 11 PM Eastern Time in Washington, DC you see the same thing as someone in Denver, CO will see at 11 PM Mountain time (i.e. 2 hours later)
- different latitudes see different parts of the sky, or the same things but at different altitudes at the same time

Local Solar Time

- when Sun is "on" the meridian, we call it "noon"
- one rotation (one day) corresponds to time it takes Sun to go all the way around and appear again on the meridian
- altitude of Sun increases from sunrise to noon (AM) and decreases from noon to sunset (PM)
- midnight corresponds to Sun on the meridian on the other side of the Earth ($1/2$ rotation or 12 hours after noon)
- in astronomy, we use local solar time, but it's not very practical for society...

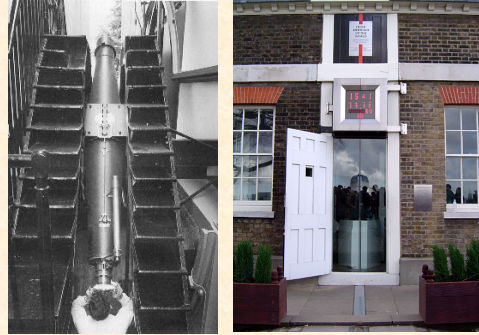
Solar v. Civil Time

- Sun "on" meridian at *local solar noon*
 - only one longitude on Earth has solar noon at any given instant
- It would get very confusing if everybody had *different* clocks, but
 - it would be just as confusing if we all used the *same* time: it could be dark at "noon"!!
 - but astronomers, pilots, etc. use "Universal Time"
- So we split the difference and have 24 time zones, each about 15° wide ($360^\circ/24h=15^\circ/h$)
 - Sun (and stars) move 15° each hour = 1 hour of RA
 - "Civil Time" = "Solar Time" $\pm 1/2$ hour

Coordinate Systems & Diurnal Motion

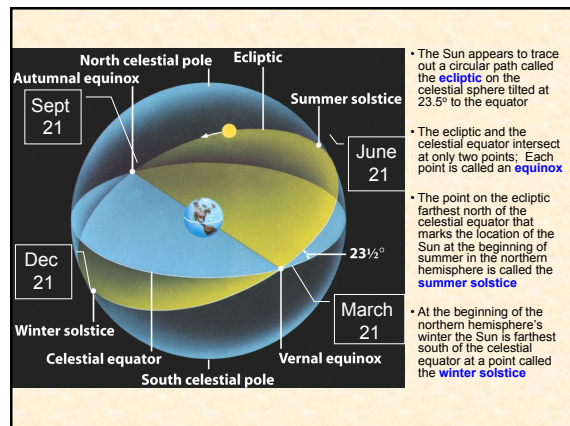
- Review:
 - Altitude, Azimuth, Zenith, Nadir, Meridian
 - Right Ascension, Declination
 - How are all of these measured? Units?
 - Apparent motion on sky depends on (1) object's declination and (2) your latitude
- Transit Altitude --> Declination
 - how do we use transits to measure RA?
- Thought experiment: Transit Altitude(s) of a Circumpolar Star
 - upper and lower culmination
 - can find meridian, NCP, equator, declination, latitude

Transit Telescope at Royal Observatory in Greenwich and the "Prime Meridian"



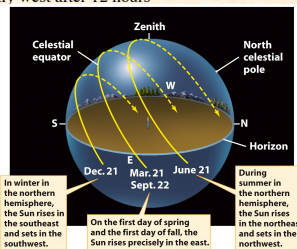
Apparent Annual Cycle of the Sun

- Sun appears to go around us once per day, but it also *appears* to go around us once per year
- Over course of year, Sun's apparent path on the sky is called the **ECLIPTIC**
- Ecliptic passes through 12 constellations of the **ZODIAC**; multiple meanings of "ecliptic"
- Sun moves 360 degrees along the ecliptic in 365 days; $\approx 1^\circ/\text{day}$; $\approx 30^\circ/\text{month}$; ≈ 2 hours of rotational motion (at $15^\circ/\text{h}$) ≈ 4 m of time
- which direction? what can you infer?



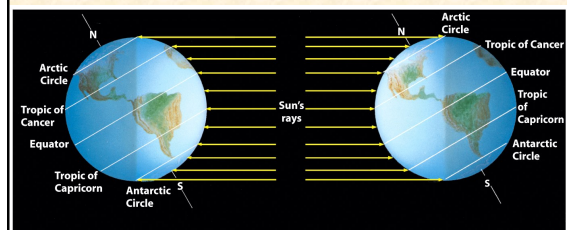
The Sun's Motion Viewed From Charleston

- **June:** rise north of east; $57+23.5=80.5$ degrees elevation on meridian at noon; set north of west > 12 hours later
- **December:** rise south of east; $57-23.5=33.5$ degrees elevation on meridian at noon; set south of west after < 12 hours
- note: difference is 47° ; no matter where you are
- **March or September:** rise exactly east; 57° above horizon on meridian at noon; set exactly west after 12 hours



Sun Viewed from The Tropics and from "The Land of the Midnight Sun"

- **Arctic Circle:** $90-23.5 = 66.5^\circ$ latitude
- **Antarctic Circle:** -66.5° latitude
- **Tropic of Cancer:** $+23.5^\circ$ latitude
- **Tropic of Capricorn:** -23.5° latitude



(a) Earth at winter solstice

(b) Earth at summer solstice

Sun's Apparent Annual Motion

- Even though Sun's motion around us is East->West, it appears to move Eastward w.r.t. the stars. $\sim 1^\circ$ or 4 min/day [demo]
- Stars rise ~ 4 min earlier each day (~ 2 h/month) [solar/civil time]
- Sun's apparent motion along ecliptic, which is tipped by 23.5° w.r.t. celestial equator, and velocity is not constant [analema:]
- Sun north of celestial equator March -> Sep
 - day > 12 h
 - meridian altitude > 57° (from Charleston)
- Sun south of celestial equator Sep -> Mar
 - day < 12 h
 - meridian altitude < 57° (from Charleston)
- Tropics: Sun at Zenith at least 1 day each year
- Arctic: Sun never rises at least 1 day each year



Anchorage Airport Sunset

How long until Sun dips below horizon?

$15^\circ/\text{h} \rightarrow 1/2^\circ$ in 2 minutes

Where on horizon will it set?

10 minutes later...

Notice how the Sun is moving almost parallel to the horizon.